

Bistable Dynamics of Turbulence Intensity in a Corrugated Temperature Profile

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- Motivations
- Mesoscale temperature profile corrugation and nonlinear drive
- Bistable spreading of the turbulence intensity:
 - subcritical excitation
 - propagation

How turbulent fluctuations penetrates stable domains?

→ Anomalous transport

→ Collapse of H-mode

Most previous works treat turbulence spreading as a Fisher front.

Conventional wisdom: Fisher front with a nonlinear diffusivity

Generic structure of Fisher spreading equation:

$$A \equiv -\partial_x T$$

$$\frac{\partial}{\partial t} I = \underbrace{\gamma_0 (\langle A \rangle - A_c) I}_{\text{linear excitation}} - \underbrace{\gamma_{nl} I^2}_{\text{nonlinear propagation}} + D_1 \frac{\partial}{\partial x} \left(I \frac{\partial}{\partial x} I \right)$$

Nontrivial solution requires:

$$\langle A \rangle > A_c \Rightarrow I \propto \langle A \rangle - A_c$$

Suffering from two serious drawbacks:

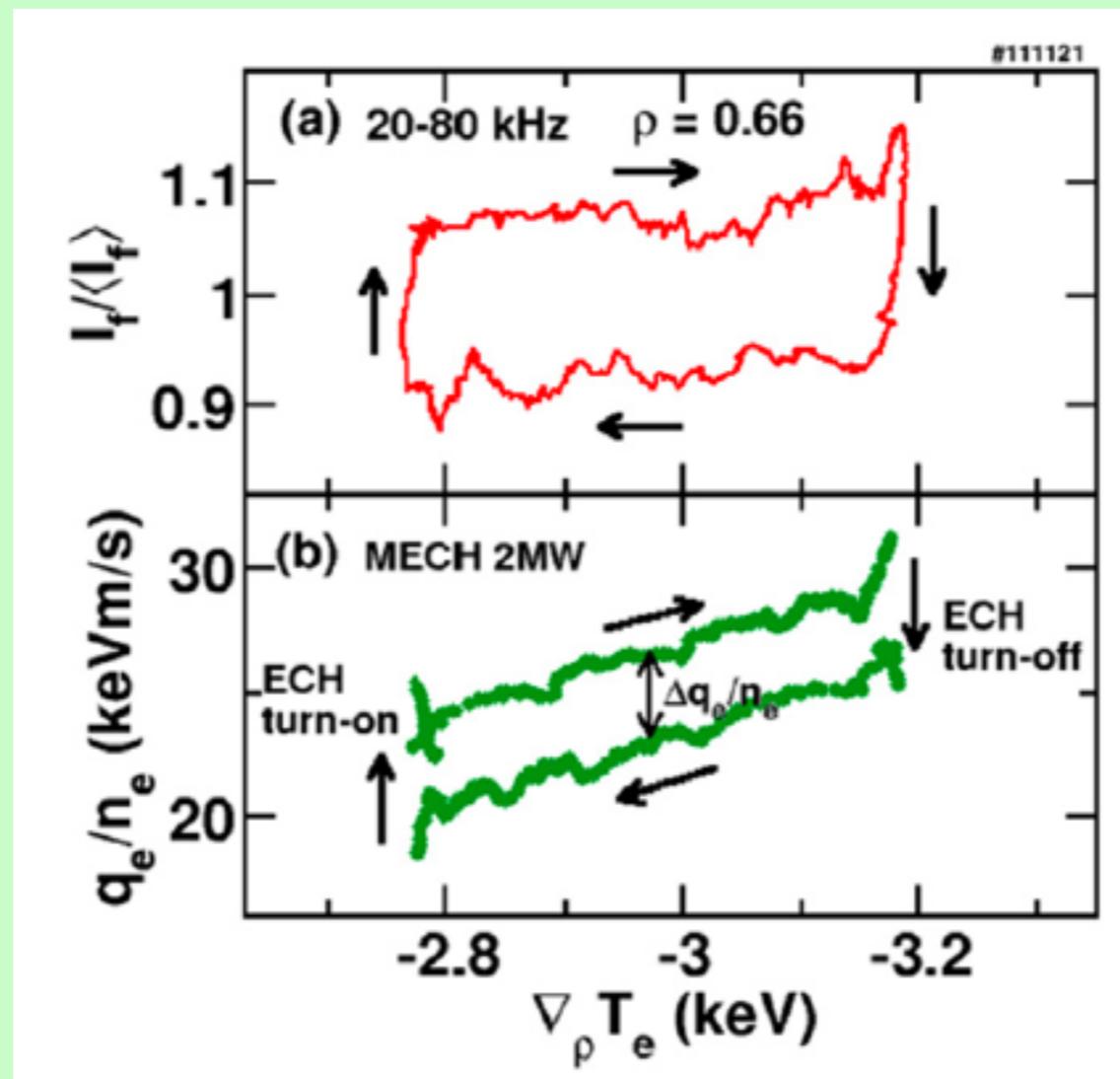
***insufficient near marginal state**

****can be strongly damped in subcritical region**

Motivations

The turbulence intensity is ***unistable*** in the Fisher model. However:

A hysteretic relation between **turbulence intensity** and temperature gradient also observed:



Inagaki2013



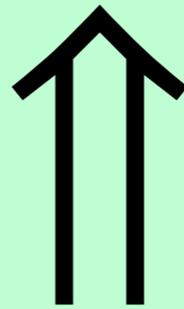
An indication of bistability of the turbulence intensity!

We missed something here??

$$\frac{\partial}{\partial t} I = \gamma_0 (\langle A \rangle - A_c + \boxed{??}) I - \gamma_{nl} I^2 + D_1 \frac{\partial}{\partial x} \left(I \frac{\partial}{\partial x} I \right)$$

In this talk, we propose the missed piece is the nonlinear drive induced by the **corrugation of the temperature field**.

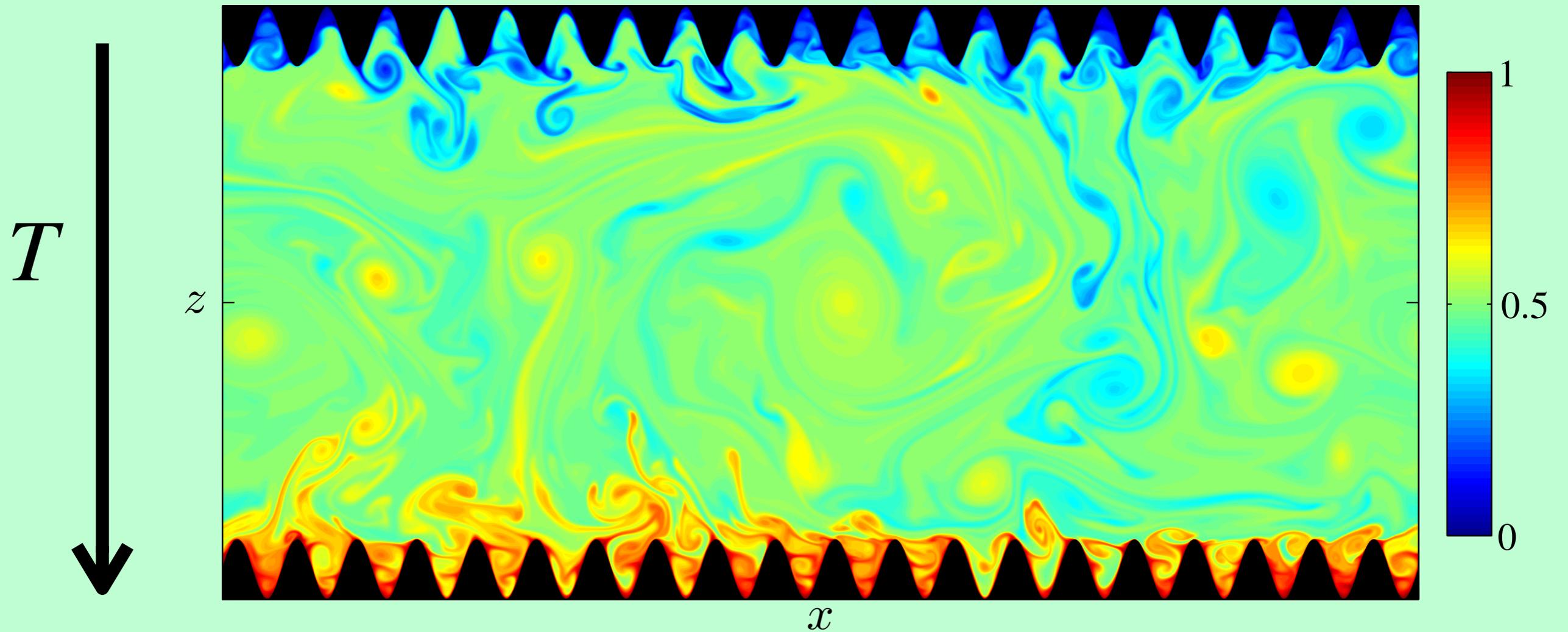
The key for *nonlinear* turbulence excitation:
temperature corrugation by inhomogeneous turbulent mixing.



Inevitable consequence of potential enstrophy conservation

A consistent treatment of
multi-scale, multi-field couplings is required...

An example from Rayleigh-Bernard convection



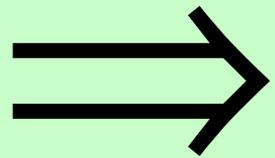
Toppaladodi et al 2017

Roughened temperature profile enhances turbulent heat flux.

How mesoscale fields impact evolution of turbulence intensity?

Drive: $(\nabla T)_{meso}$

Dissipation: $\langle V \rangle'_{ZF} \xrightarrow{\text{local force balance}} \langle V \rangle'_{ZF} \propto (\nabla^2 T)_{meso}$



Generally, drive&dissipation act in different regions.

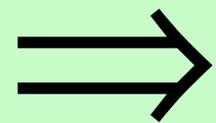
How the turbulence intensity is **excited and spreads** in the presence of a corrugated temperature profile?

The Model: bistable turbulence Intensity

The basic structure of I 's evolution is

$$\frac{\partial}{\partial t} I = \gamma_0 \left(\langle A \rangle - A_c + \underbrace{\Theta(\tilde{A}_m) \tilde{A}_m}_{\text{nonlinear drive}} \right) I - \gamma_{nl} I^2 + D_1 \frac{\partial}{\partial x} \left(I \frac{\partial}{\partial x} I \right)$$

For a mean field approximation, $\Theta(\tilde{A}_m) \tilde{A}_m = \langle \Theta(\tilde{A}_m) \tilde{A}_m \rangle + \overline{\Theta(\tilde{A}_m) \tilde{A}_m} \simeq \langle \Theta(\tilde{A}_m) \tilde{A}_m \rangle$



$$\frac{\partial}{\partial t} I = \gamma_0 \left(\langle A \rangle + \langle \Theta(\tilde{A}_m) \tilde{A}_m \rangle - A_c \right) I - \gamma_{nl} I^2 + D_1 \frac{\partial}{\partial x} \left(I \frac{\partial}{\partial x} I \right)$$

Relation between

$\langle \Theta(\tilde{A}_m) \tilde{A}_m \rangle$ and I ?

Strength of Mesoscopic ∇T Fluctuations

$$\frac{\partial}{\partial t} T + \nabla \cdot \mathbf{Q}_T = \chi_{neo} \frac{\partial^2}{\partial x^2} T + S \delta(x) \quad (*)$$

$$T = \langle T \rangle + \tilde{T} = \langle T \rangle + \tilde{T}_m + \tilde{T}_s, \quad \mathbf{Q}_T = \tilde{\nu} T, \quad \tilde{T}_m : \text{meso scale}; \quad \tilde{T}_s : \text{micro scale}$$

Define two types of average:

$$\langle \dots \rangle_s - \text{micro timescale}; \quad \langle \dots \rangle_m - \text{meso timescale} \quad \Rightarrow \quad \langle \langle T \rangle_s \rangle_m = \langle T \rangle_m \equiv \langle T \rangle$$

Multiplying T on both sides of (*) and carrying out a double average $\langle \langle \dots \rangle_s \rangle_m$ yields

Entropy balance of the turbulence:

$$A_0 \langle \mathbf{Q}_T \rangle_m + \langle \langle \tilde{A}_m \tilde{\nu} \tilde{T} \rangle_s \rangle_m \approx \chi_{neo} \langle \tilde{A}_m^2 \rangle_m$$

entropy production due to turbulent mixing of the mean temperature field

triple coupling between micro and meso scales

entropy dissipation due to neoclassical diffusion

The essential process for subcritical turbulence excitation

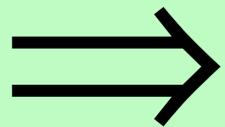
A closure on the triple coupling: 'negative' thermal conductivity

$$\langle \tilde{A}_m \tilde{v} \tilde{T} \rangle_s = \tilde{A}_m \langle \tilde{v} \tilde{T} \rangle_s$$

up gradient heat flux on mesoscale $\langle \tilde{v} \tilde{T} \rangle_s = \chi_m \tilde{A}_m = -|\chi_m| \tilde{A}_m$

$\chi_m < 0$ the negative diffusivity.

The underlying physics: roll-over of Q_m vs $\partial_x T_m$ duo to ZF shear.

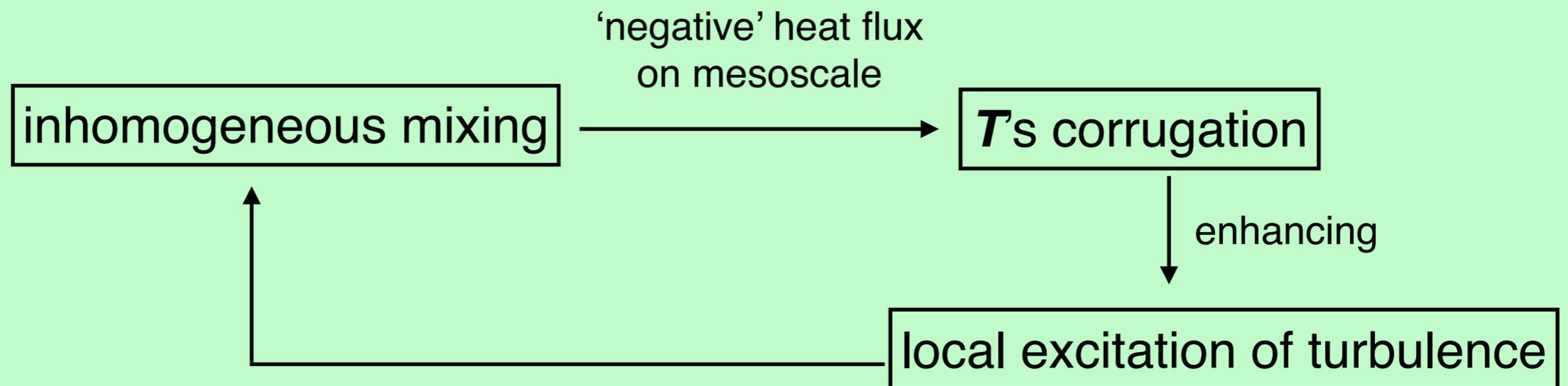


Zeldovich relation in multi-scale coupling system:

$$\langle \tilde{A}_m^2 \rangle = \frac{D_0 \langle A \rangle^2}{\chi_{neo} + |\chi_m|} I$$

T 's profile gets corrugated by the inhomogeneous turbulent mixing.

The closed loop



Bistable spreading of the turbulence intensity: subcritical excitation

I 's evolution with subcritical drive (**Fitzhugh-Nagumo type, not Fisher!**)

$$\frac{\partial I}{\partial t} = \gamma_0 (\langle A \rangle - A_C) I + \sqrt{\frac{\gamma_0^2 D_0 \langle A \rangle^2}{\chi_{neo} + |\chi_m|}} I^{3/2} - \gamma_{nl} I^2 + D_1 \frac{\partial}{\partial x} \left(I \frac{\partial I}{\partial x} \right)$$

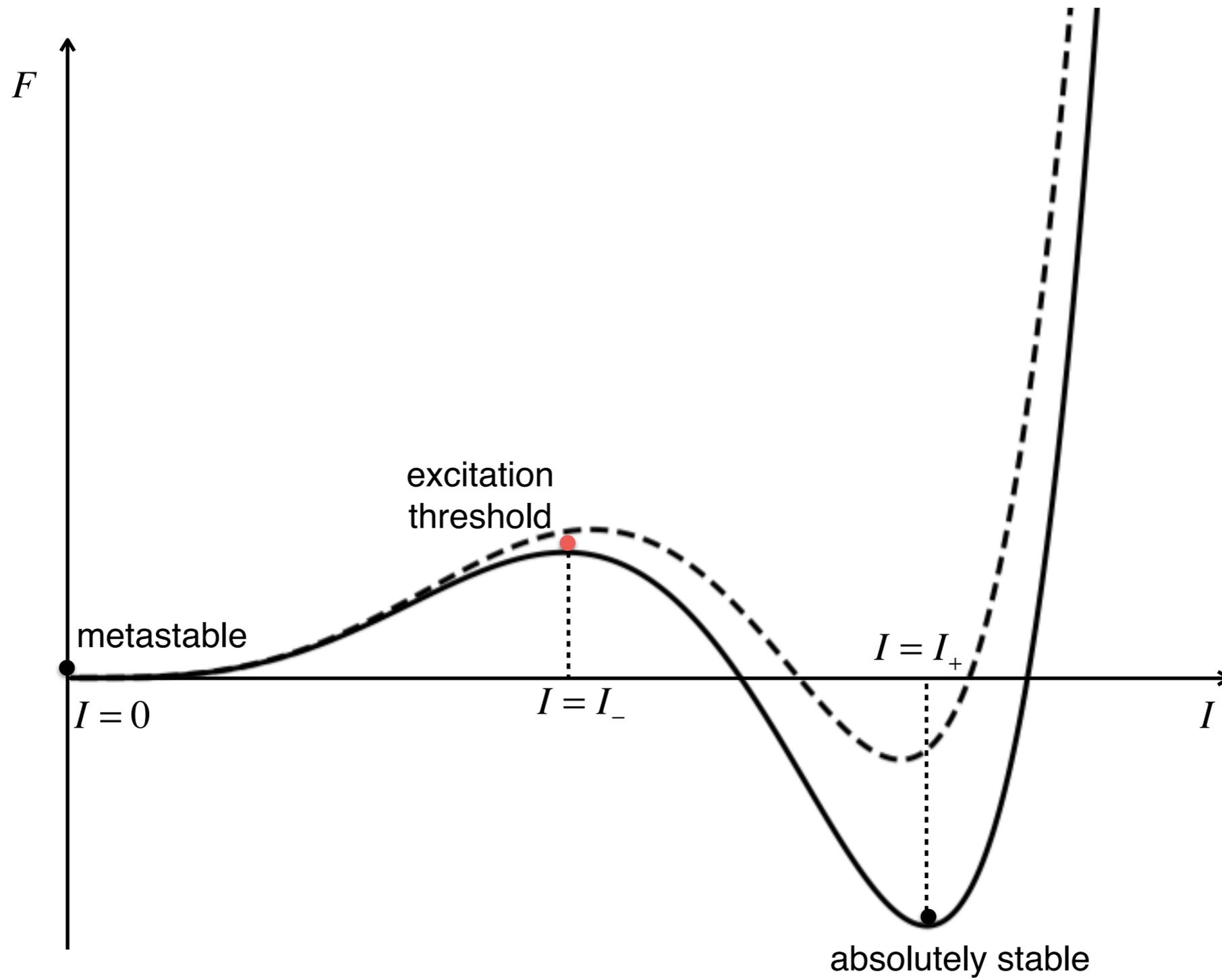
On the subcritical excitation

$$\frac{\partial I}{\partial t} = -\frac{\delta F(I)}{\delta I}$$

$$F(I) = -\frac{1}{2} \gamma_0 (\langle A \rangle - A_C) I^2 - \frac{2}{5} \sqrt{\frac{\gamma_0^2 D_0 \langle A \rangle^2}{\chi_{neo} + |\chi_m|}} I^{5/2} + \frac{1}{3} \gamma_{nl} I^3$$

Two stable solutions: $I = 0$ and $I = I_+$

One unstable solution: $I = I_-$



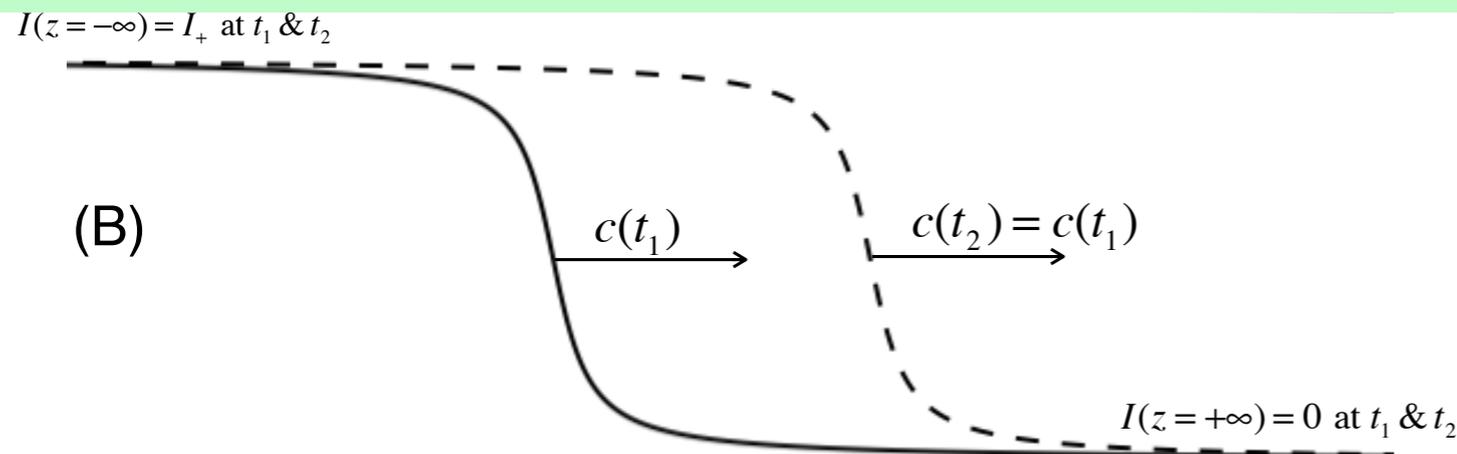
How the bistable spreading happens?

Looking for wave-like solution: $I(x,t) = I(z)$ with $z = x - c^*t$

$$-c^* \frac{d}{dz} I = -\frac{\delta F}{\delta I} + D_1 I \frac{d^2}{dz^2} I + D_1 \left(\frac{d}{dz} I \right)^2$$

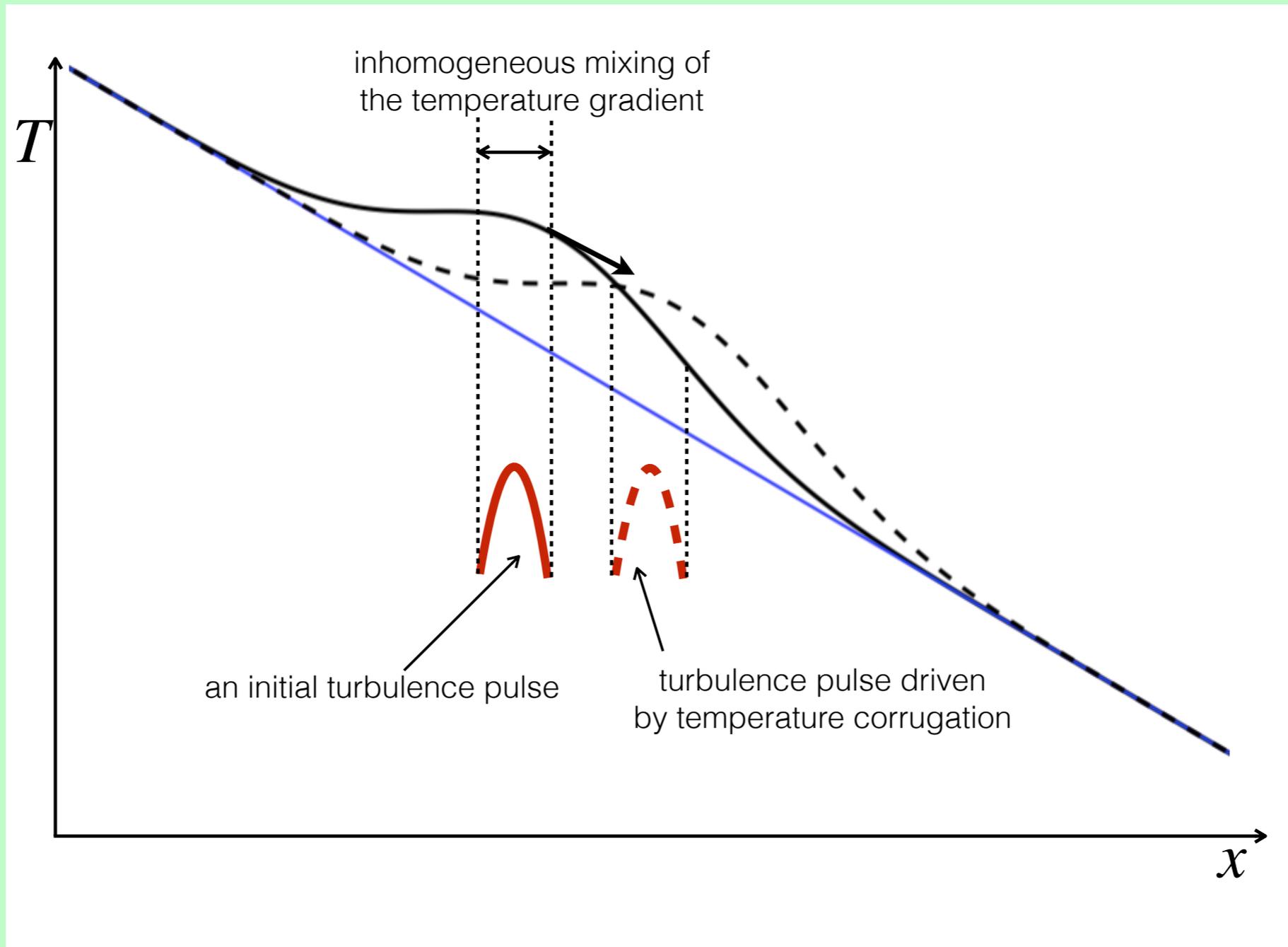
Propagation speed of the bistable front:

$$c^* = \frac{F(+\infty) - F(-\infty)}{\int_{-\infty}^{+\infty} I'^2 dz} + \frac{-\frac{D_1}{2} \int_{-\infty}^{+\infty} I'^3 dz}{\int_{-\infty}^{+\infty} I'^2 dz}$$



$$: C^* > 0$$

How the bistable spreading happens?



Summary&future work

Mesoscale temperature profile corrugations provide a natural way for subcritical turbulence excitation, and the following spreading.

Next:

Temperature profile evolution needs to be treated in a more consistent way;

Stability problem of the front, i.e., can the front be splitter by any external/internal noise?